

# Hamiltonian analysis of Holst action

Consider a generalization of the Palatini action for classical General Relativity with a topological term — the *Holst action*:

$$S[\underline{e}, \underline{\omega}] = \frac{1}{16\pi G} \int \underline{e}^I \wedge \underline{e}^J \wedge \left( \underline{*F}_{IJ} + \frac{1}{\gamma} \underline{F}_{IJ} \right) = \frac{1}{16\pi G} \int e \cdot e_I^\alpha e_J^\beta \left( F_{\alpha\beta}{}^{IJ} - \frac{1}{2\gamma} \epsilon^{IJ}{}_{KL} F_{\alpha\beta}{}^{KL} \right),$$

where the Cartan structure equations define the torsion and curvature as

$$\begin{aligned} \underline{T}^I &= \underline{\nabla} \underline{e}^I = \underline{d} \underline{e}^I + \underline{\omega}^I{}_J \wedge \underline{e}^J, \\ \underline{F}^I{}_J &= \underline{d} \underline{\omega}^I{}_J + \underline{\omega}^I{}_K \wedge \underline{\omega}^K{}_J, \end{aligned}$$

and the Hodge star acts on flat indices as

$$\underline{*F}_{IJ} = \frac{1}{2} \epsilon_{IJKL} \underline{F}^{KL}.$$

## Canonical variables

After the spacetime foliation, we proceed with an unusual choice of canonical variables — the *Ashtekar-Barbero connection*:

$$\begin{cases} A_a^i = \frac{1}{2} \epsilon^i{}_{jk} \omega_a^{jk} + \gamma \omega_a^{0i}, \\ E_i^a = \frac{1}{2} \epsilon_{ijk} \epsilon_{abc} e_b^j e_c^k. \end{cases}$$

Note that while  $A_a^i$  is a true 3-dimensional vector,  $E_i^a$  is actually a vector density, since it contains the densitized Levi-Civita symbol  $\epsilon_{abc}$  in its definition.

Also,  $A_a^i$  does not have nice  $SO(3,1)$  transformation properties, unless a specific value is given to the *Barbero-Immirzi constant*:  $\gamma = i$ , in which case we recover the usual *Ashtekar connection* which transforms in the self-dual representation of  $SO(3,1)$ . But unless  $\gamma$  is real, the Ashtekar-Barbero connection becomes complex, and nontrivial reality conditions have to be imposed afterwards.

Nevertheless, objects  $A_a^i$  and  $E_i^a$  are well-defined in all gauges and coordinate systems, despite having nonlinearities in their transformation properties.

The vector density  $E_i^a$  has a nice geometrical interpretation. Consider a surface embedded into space:  $x^a = \chi^a(\sigma^1, \sigma^2)$ . The area of the surface is given by an integral

$$A = \int d^2\sigma \sqrt{\det f}$$

with the induced metric given by

$$f_{xy} = q_{ab}(\chi) \cdot \frac{\partial \chi^a}{\partial \sigma^x} \frac{\partial \chi^b}{\partial \sigma^y} = \delta_{ij} e_a^i(\chi) e_b^j(\chi) \frac{\partial \chi^a}{\partial \sigma^x} \frac{\partial \chi^b}{\partial \sigma^y},$$

It follows that

$$A = \int d^2\sigma \sqrt{E_i^a n_a E_i^b n_b},$$

where

$$n_a(\sigma) = \epsilon_{abc} \frac{\partial \chi^b}{\partial \sigma^1} \frac{\partial \chi^c}{\partial \sigma^2}.$$

Thus,  $E_i^a$  contains information about areas of surfaces, and indeed, after appropriate regularization, from this classical quantity the quantum area operator of LQG is built.

## Hamiltonian analysis

Now we have to rewrite the Holst action in terms of the Ashtekar-Barbero connection.

$$F^{IJ}{}_{\mu\nu} = \partial_{[\mu} \omega_{\nu]}{}^{IJ} + \eta_{KL} \omega_{[\mu}{}^{IK} \omega_{\nu]}{}^{LJ}$$